



# Cosmic Rays

What are they?

How do we see the invisible?

10 muons per second go through you!  
You can't feel them but can detect  
them!

Mark Adams  
QuarkNet at Fermilab



# QuarkNet Detectors

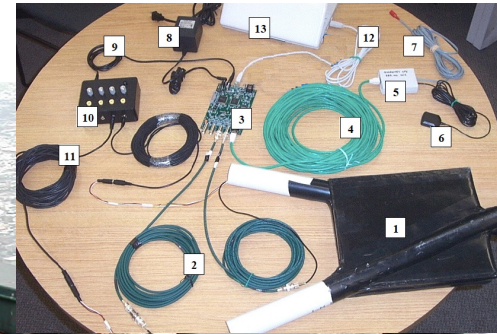
## High Energy Physics technology provided to high schools



Build telescope



Roof of Sears Tower





# Cosmic Ray Discovery

- Cosmic Rays discovered in 1912 (Victor Hess) in balloon experiments
- Radiation higher at 5000m than at sea level – implies source hitting atmosphere
- No difference during partial eclipse – implies Sun not the source



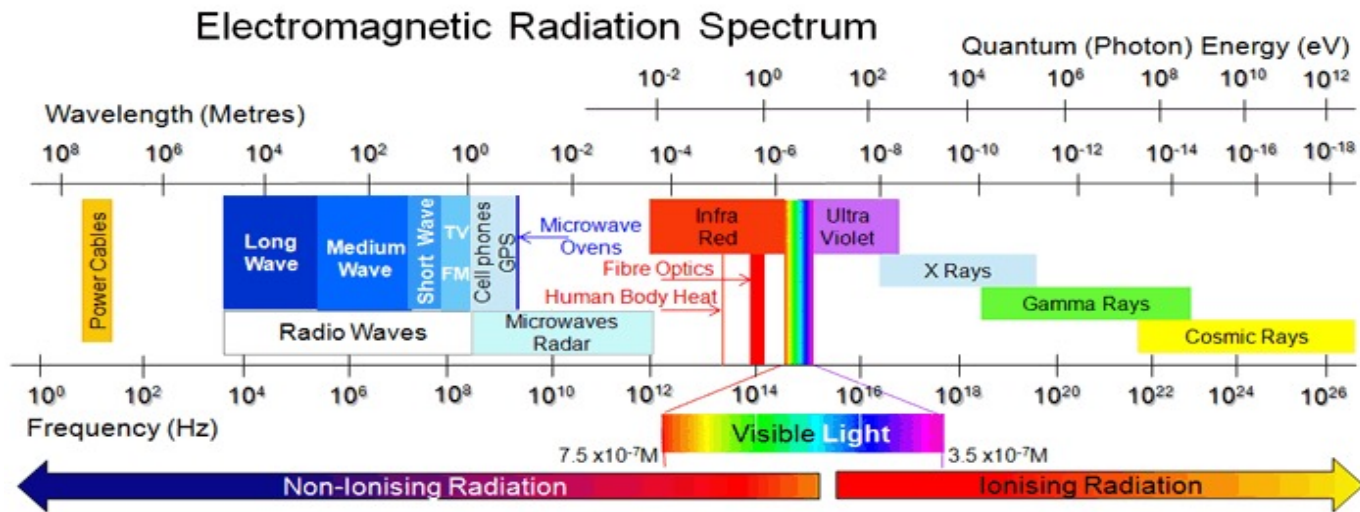


- Really!?! Cosmic rays we see on earth don't come from the Sun?
- Why?
- They are particles with energies much higher than the energy available in the sun



# How High is High Energy?

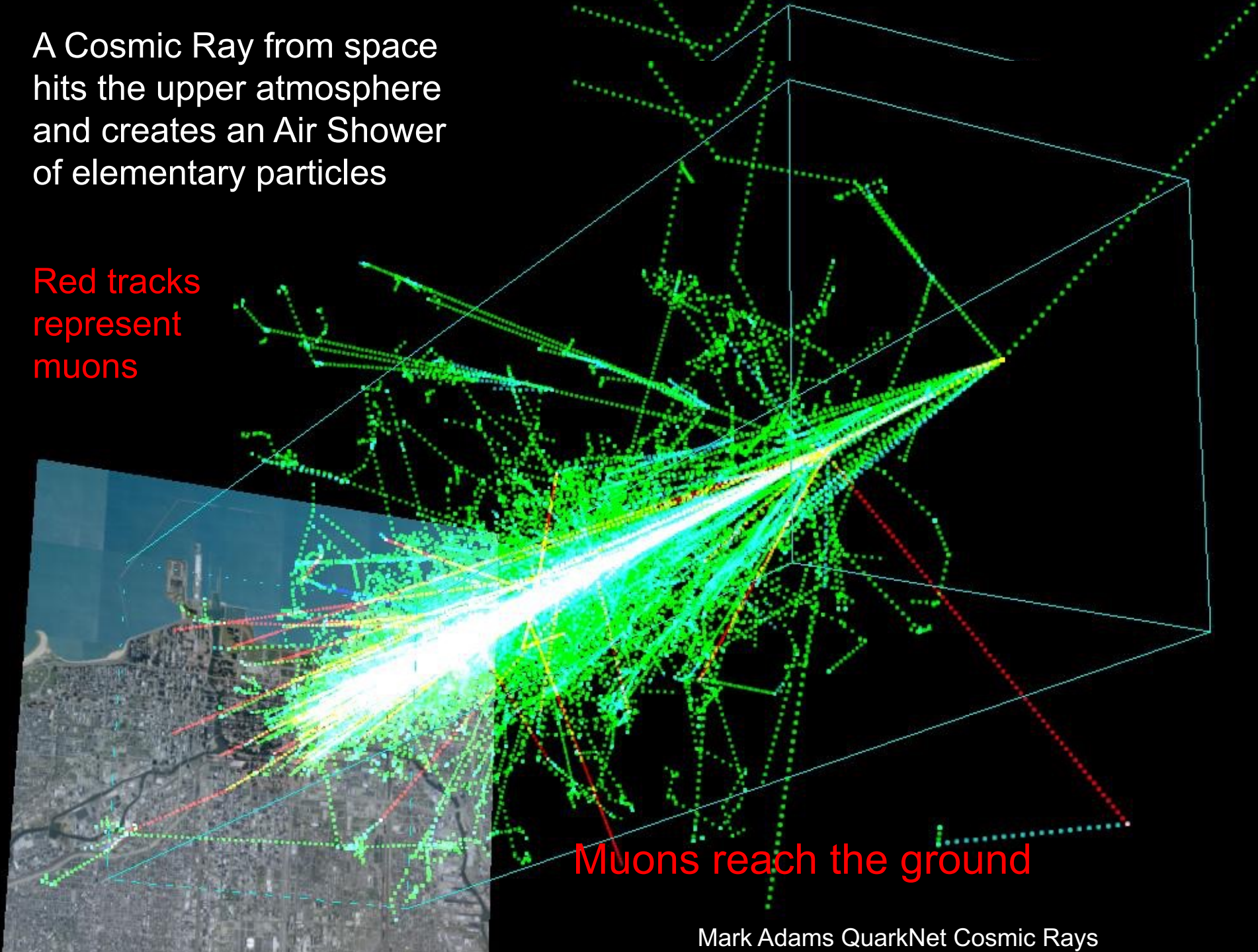
- 1 eV Sun light photons - atomic energies
- 1000 eV x-rays (keV)
- 1,000,000 eV nuclear physics in sun's core (MeV)
- 1,000,000,000 eV gamma rays (GeV)



Photons used here for energy example – but cosmic rays are charged and  $E > \text{GeV}$

A Cosmic Ray from space hits the upper atmosphere and creates an Air Shower of elementary particles

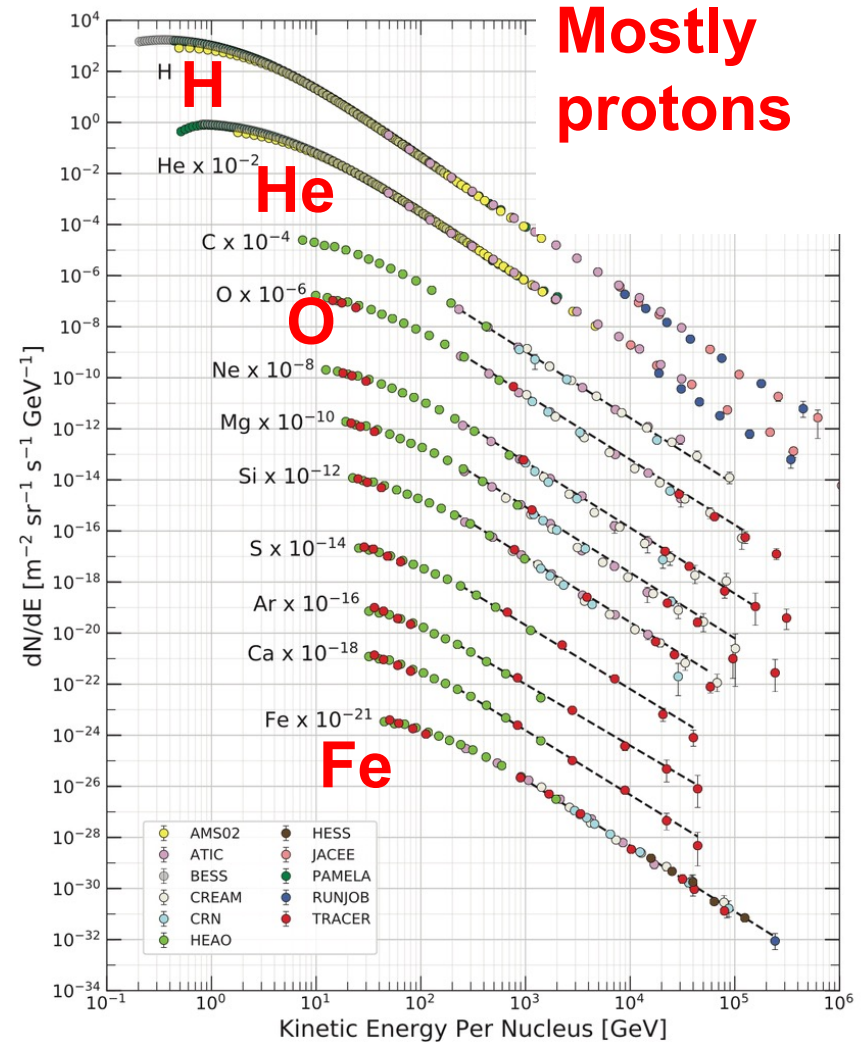
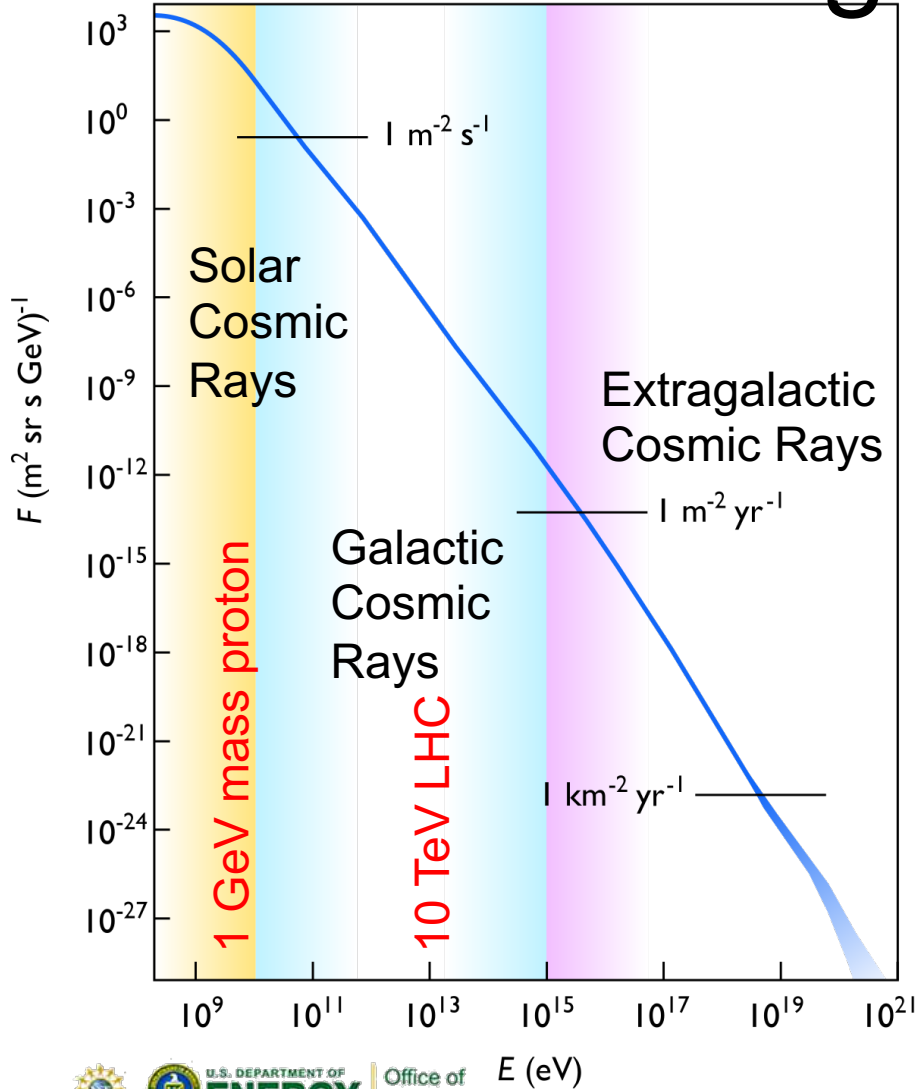
Red tracks represent muons



Muons reach the ground



# Cosmic Ray Rates Hitting the Atmosphere





# Cosmic Ray Sources

Galactic - Supernovae



Extragalactic – Active Galaxies



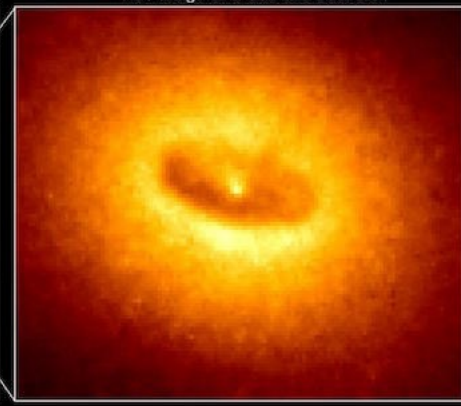
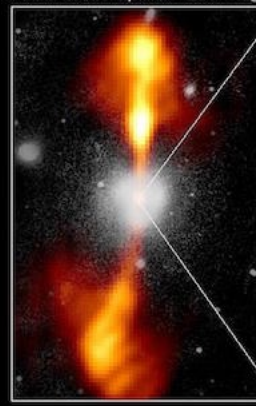
Natural very high energy accelerators

Core of Galaxy NGC 4261

Hubble Space Telescope  
Wide Field / Planetary Camera

Ground-Based Optical/Radio Image

HST Image of a Gas and Dust Disk



380 Arc Seconds  
88,000 LIGHTYEARS

17 Arc Seconds  
400 LIGHTYEARS



# Source of Cosmic Rays

**Primary Cosmic Rays**

**Death of massive star**

Nuclei fusion in stars

Supernova explosion

Proton

Helium

Carbon

Oxygen

Silicon

Iron

Primary elements, protons, He, C, O, Ne, Mg, Si..., Fe nuclei are produced during the lifetime of stars.

They are accelerated in supernovae explosions and expelled in the interstellar medium where they propagate diffusively through the galaxy.

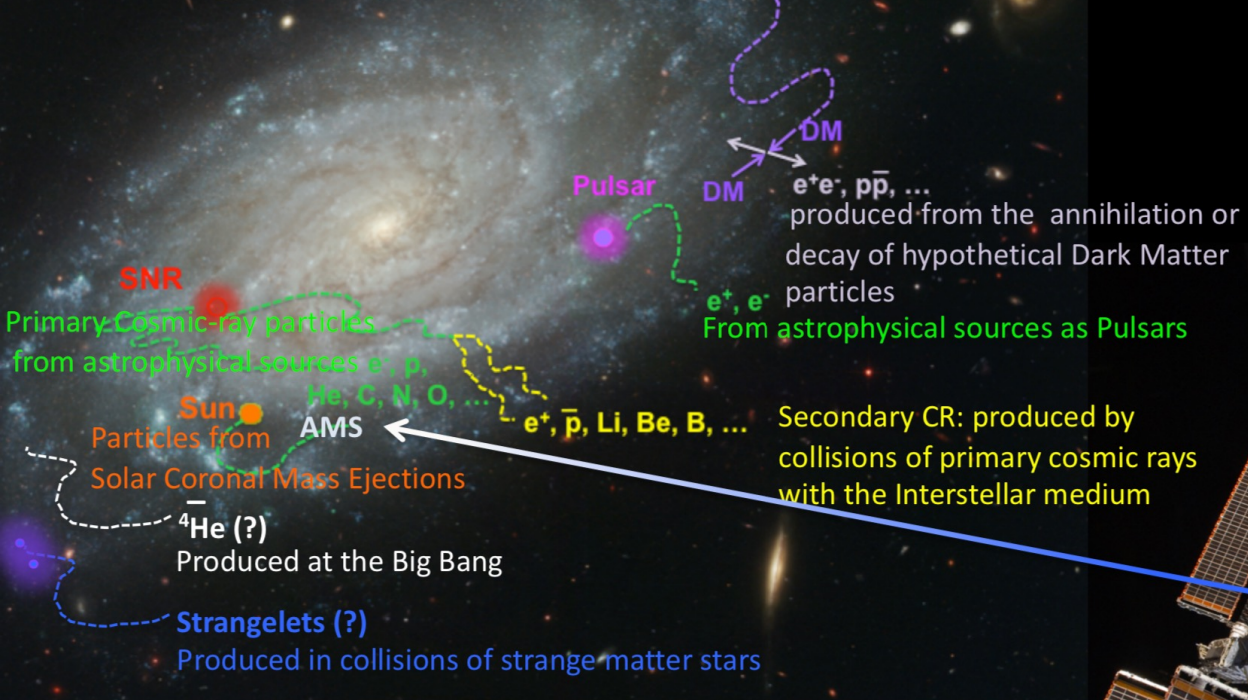
*Slide from Dr. Mercedes Paniccia AMS (Alpha Magnetic Spectrometer) talk at CERN*



# Cosmic Rays before hitting atmosphere

## Charged particles of Cosmic Origin

**Mostly protons  
Curvy paths  
through galaxy**



**AMS-02 webpage:**  
<https://ams02.space>

*Slide from Dr. Mercedes Paniccia talk at CERN*



# Cosmic Ray Summary

Cosmic Rays are high energy nuclei (mostly proton) accelerated in some extreme condition – eventually hitting Earth’s atmosphere

Galactic – Exploding massive stars

Extragalactic – Massive Black Hole accelerates particles

Weird that source of this energy is Gravity

**Charged – trajectory bent by magnetic fields**

Can’t point back to their origin

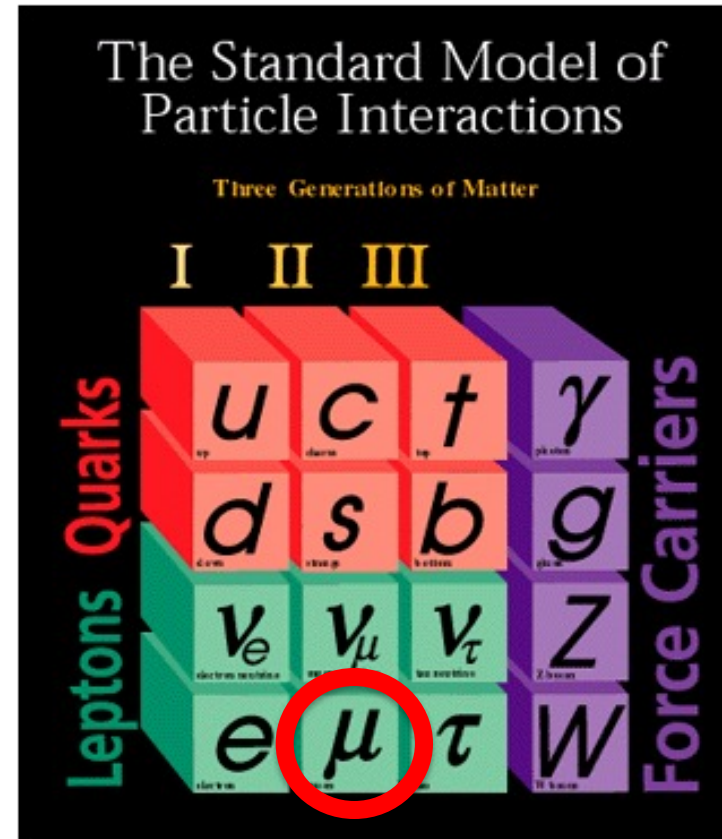


- Cosmic Ray Shower Video
- <https://vms.fnal.gov/asset/detail?recid=1963802>



# Muons?

- Muons are created in the cosmic ray interactions with the atmosphere
- Particles that reach ground have large energies (2GeV)
- Earth's magnetic field shields us from most Cosmic Rays





# Air Showers make muons

- Many particles created in collisions with atmosphere (Air Shower) ~ 15 km above surface
- Most particles interact strongly and stop
- some decay (to muons); muons live long enough to reach surface
- Muon lifetime is ~2 microseconds – at speed of light that is 600m – so, again, why do they reach the ground?
- Why do muons make it to ground?

(demonstrates Einstein's Special Relativity)



# Muons at the surface

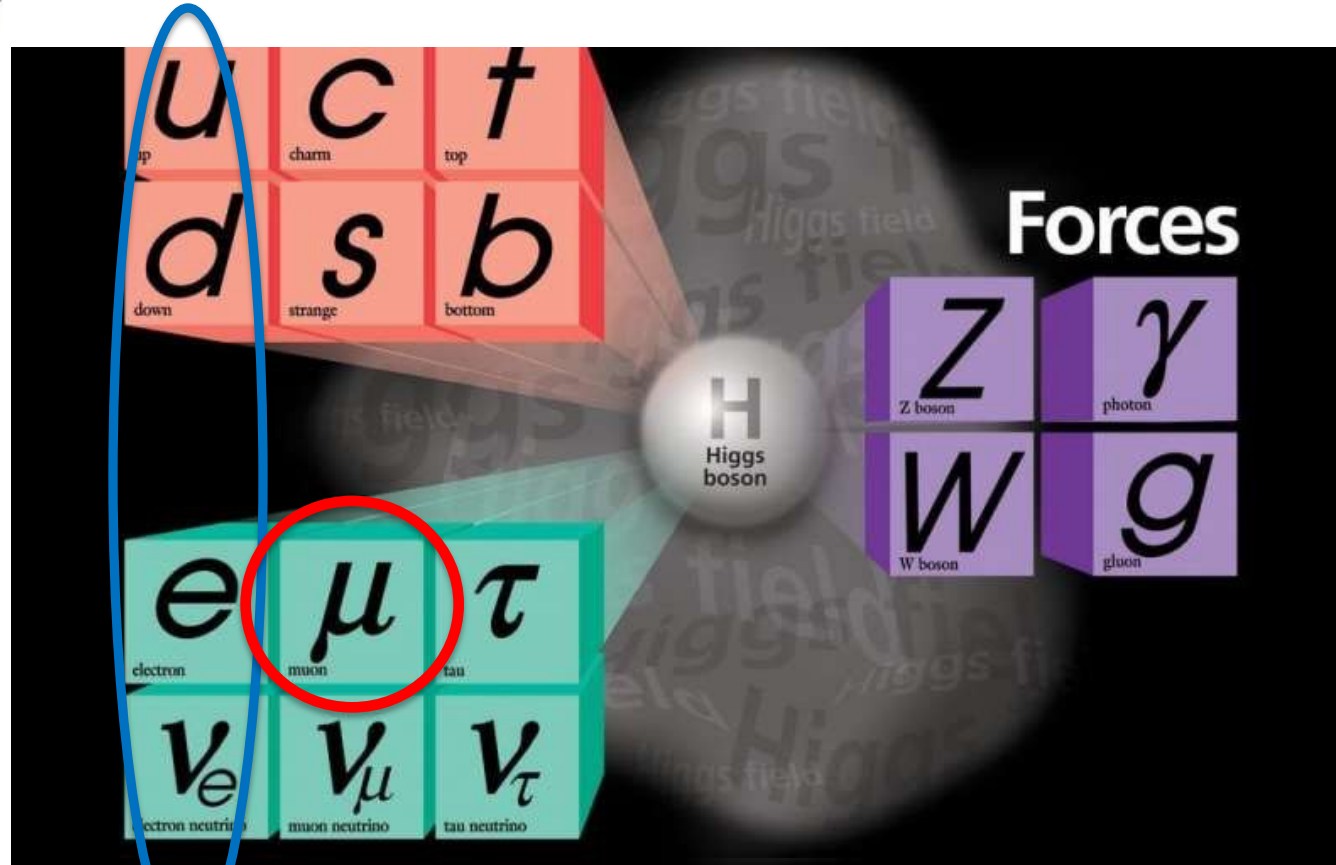
- Muons discovered (identified in air showers) in 1936 (Anderson and Neddermeyer). Rabi “Who ordered that?”
- Fundamental particle: charged, unstable lepton like a heavy (0.105 GeV) electron
- Particle sweet spot –
  - does notice nuclei (lepton); Loses energy electromagnetically through atmosphere gradually
  - **Lives long enough to reach surface**
  - Massive (unlike electron) - travels far



# The Standard Model

quarks make up protons and neutrons using Strong force

leptons don't feel the strong force



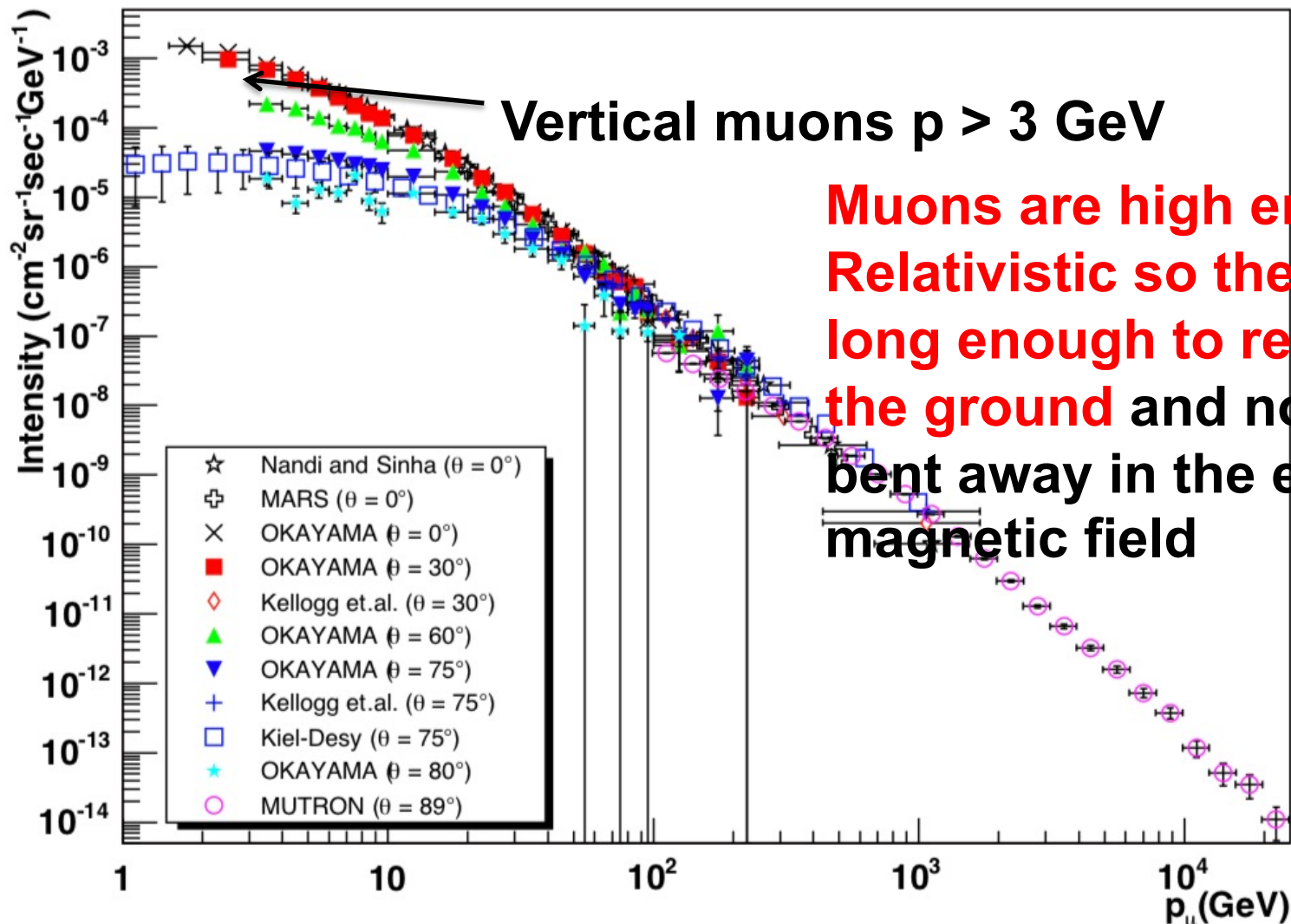
Our world

Fermions take up space

Bosons are carriers of force



# Muon momentum at surface



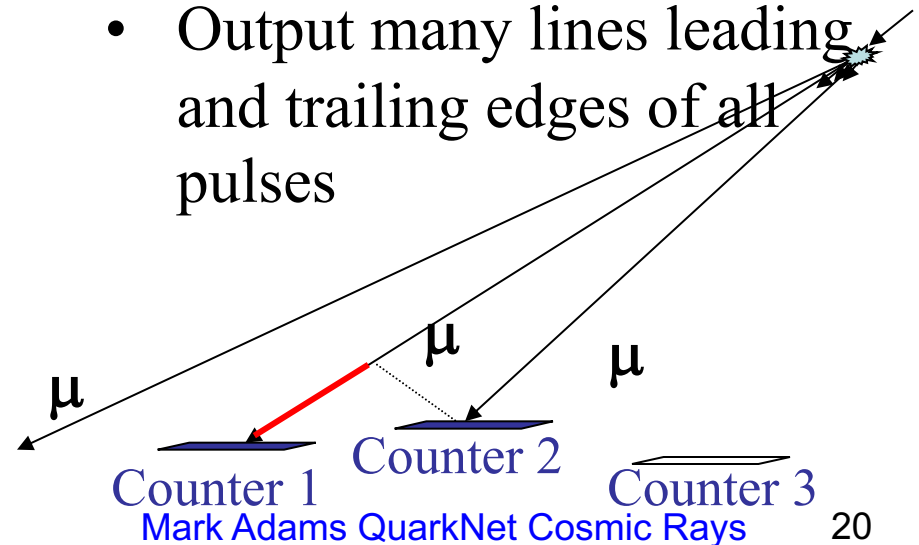
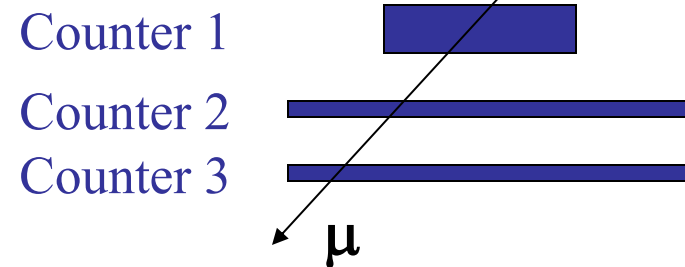
$10^9$  eV



# Compare single and multiple muon operation

- Time Dilation; Muon Lifetime
- Single muon (plus decay)
- Energy – 1 GeV
- Trigger 3-fold vertical array
- Measure muon rates and muon decay time
- Speed with TOF

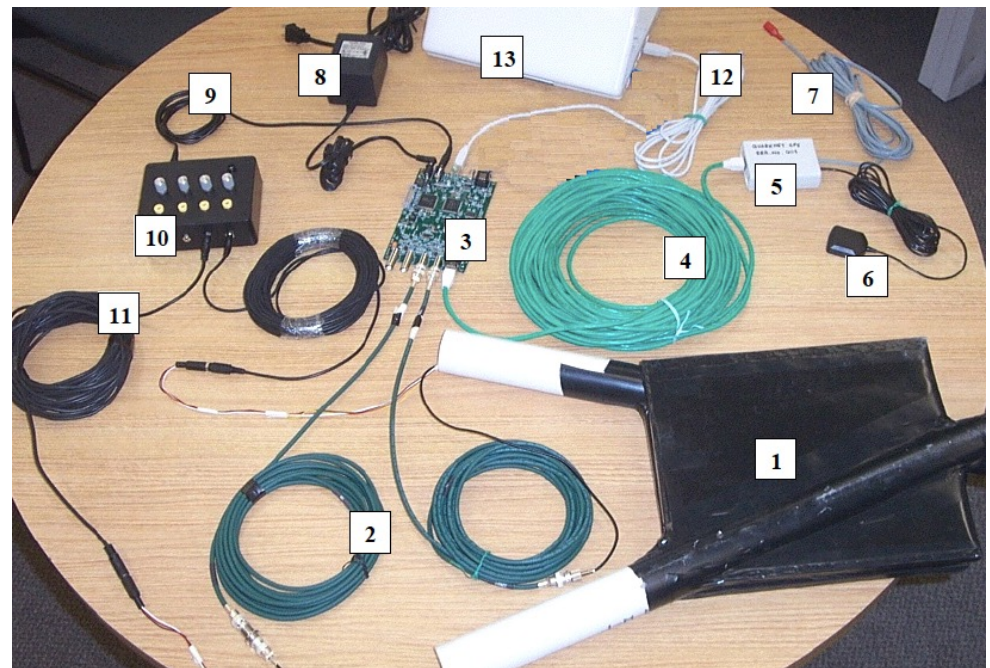
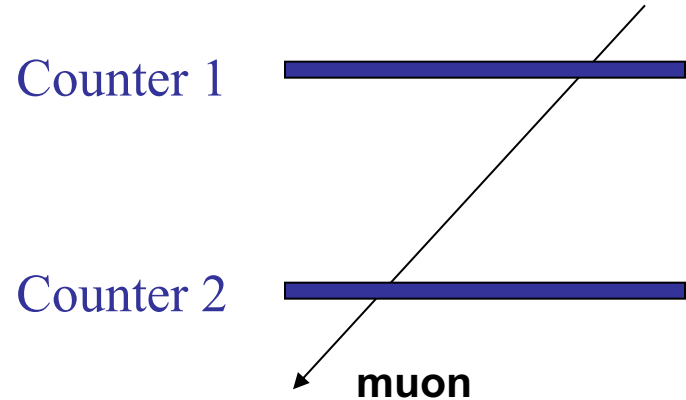
- Large Array (DAQII)
- Multiple muons - air shower
- Energy - 100 GeV
- Trigger 2 in horizontal array
- Measure rates and pattern - correlate with other arrays
- Output many lines leading and trailing edges of all pulses





# QuarkNet Detector

- Single muon passes through all detectors
- Energy  $> 2$  GeV
- Electronics selects events with 2 hit counters
- GPS gives absolute time
- A detector stack can point
- Measure Muon Speed with Time of Flight (TOF)





# What can we measure?

Muons tell us about the cosmic rays, but we also can study the muons themselves

- Prove that our counters detect muons and not noise
- What direction do muons come from?
- Rates of muons versus zenith direction
- Rates of muons versus separation
- **Speed of muons**
- **Muon Lifetime**
- **Characteristics of Air Shower**
- Muon g-2; Moon Shadow; Pyramid (like your dome)



# Conclusions

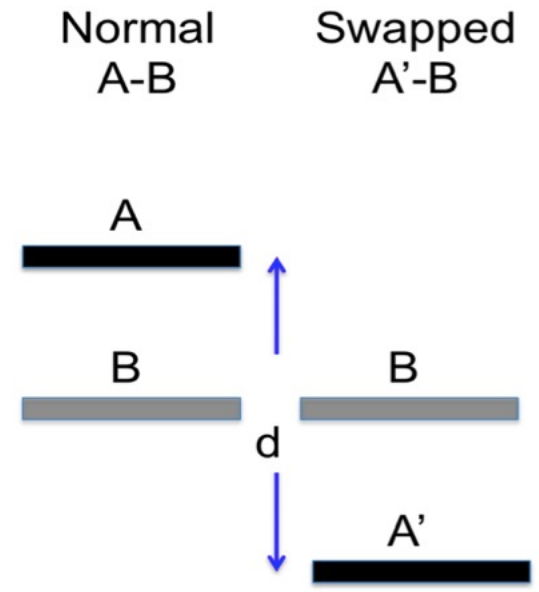
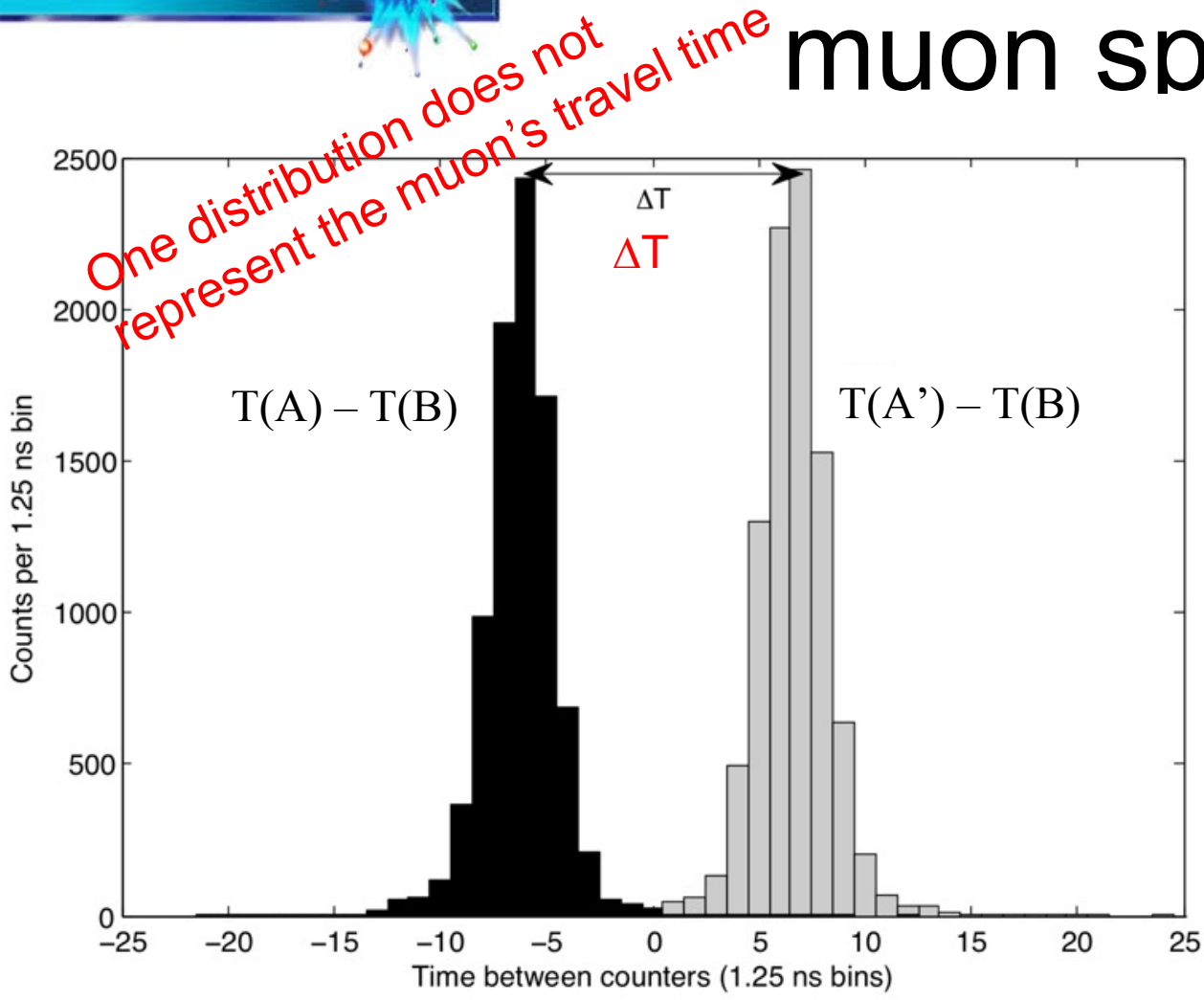
- Cosmic Rays rain down on the Earth from supernovae and Active Galaxies
- High schools use QuarkNet detectors to study cosmic rays and their resulting muons
- 100,000 files of data are available in e-Lab for all to study
- I hope you enjoy proving muons are real
- May measure maximum speed in the Universe this week



# Extra Slides

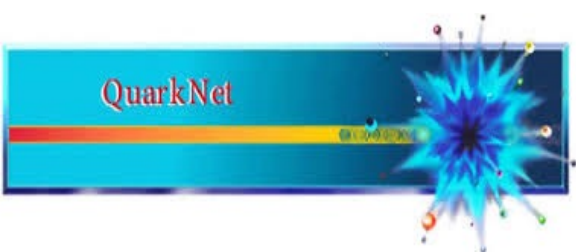


# Timing Distributions muon speed



**Speed =  $\Delta z / \Delta T$**

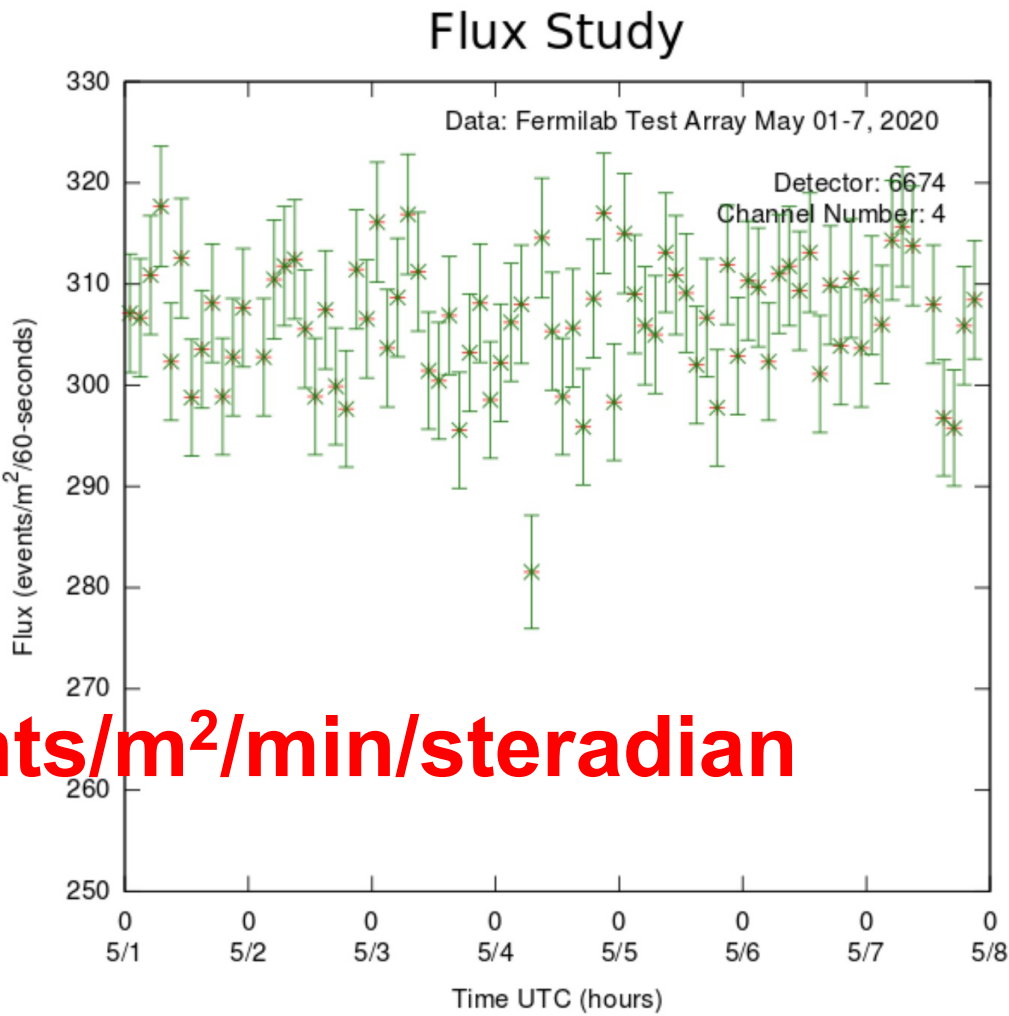
GBS HS students published a measurement of the muon's speed



# Flux - Rate versus time

**A week of data**  
**Note: no big day/night effect**  
**> Not from Sun**

**Number of events/m<sup>2</sup>/min/steradian  
in 6-hr bins**





# Shower from Dense Array

